

# Testing formulas for decay rates

## Equations for decay rates

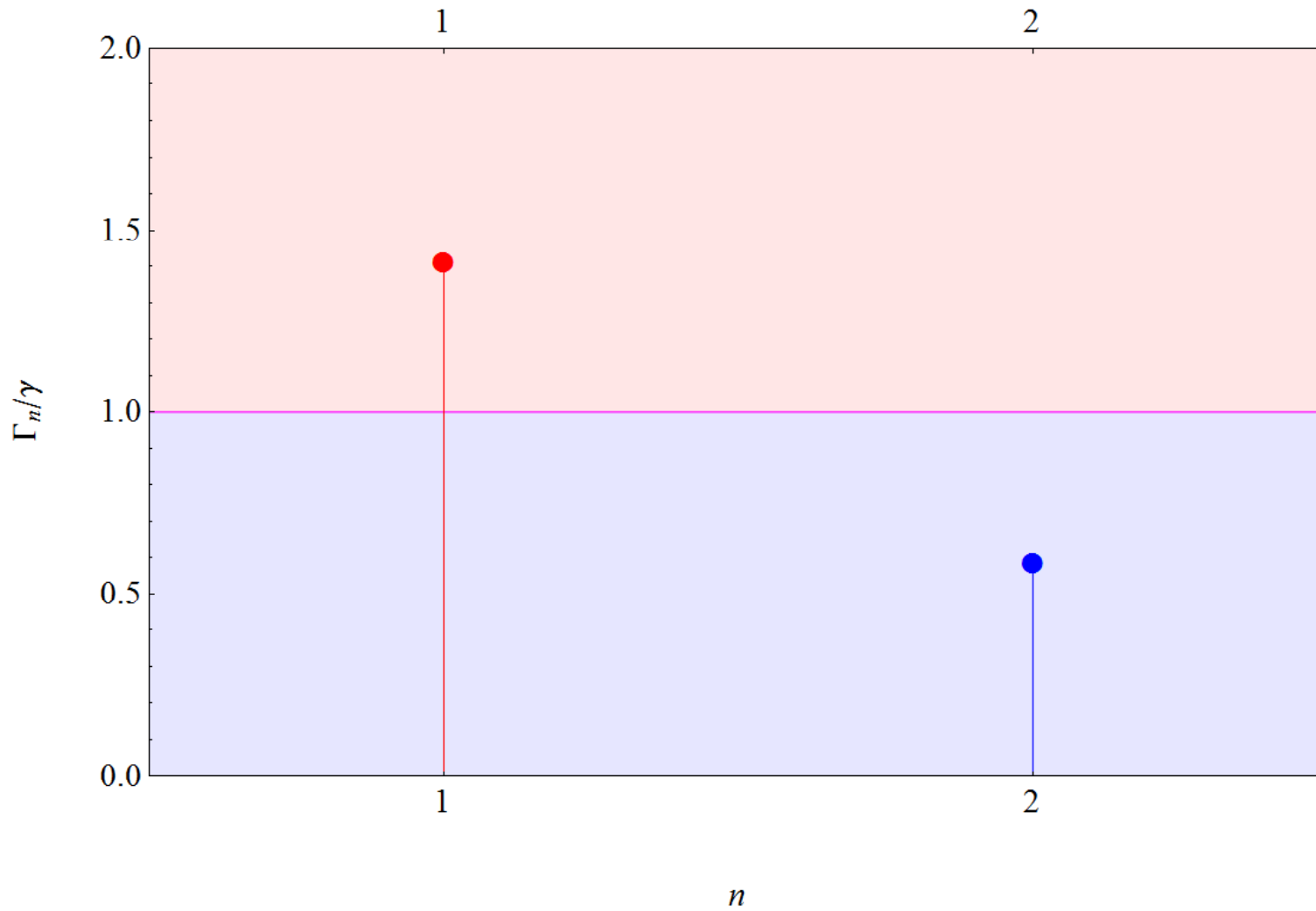
$$\left\{ \Gamma_i^{(N)} /_{i=1,2,\dots,N} \right\} = \gamma \times \text{eigenvalues}(\mathbf{F})$$

$$F_{ij} = \begin{cases} 1, & i = j \\ \frac{\sin(k_0 |\vec{r}_i - \vec{r}_j|)}{k_0 |\vec{r}_i - \vec{r}_j|}, & i \neq j \end{cases}$$

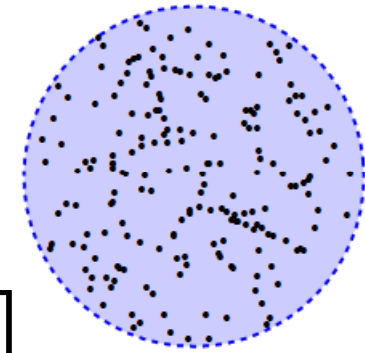
$$N = 1: \quad \Gamma_1^{(N)} = \gamma$$

$$N = 2: \quad \Gamma_1^{(N)} = \gamma(1 + F_{12}), \quad \Gamma_2^{(N)} = \gamma(1 - F_{12})$$

$$\lambda/R_{12} = 3., \quad N = 2$$



$N$  atoms inside a sphere:  
theoretical prediction



$$\Gamma_i^{(N)} \approx \frac{3N}{2} \left[ j_n^2(k_0 r) - j_{n-1}(k_0 r) j_{n+1}(k_0 r) \right],$$

where  $n = \{i^{1/2}\} - 1$

## Coherent Emission of a Photon by Many Atoms

V. ERNST

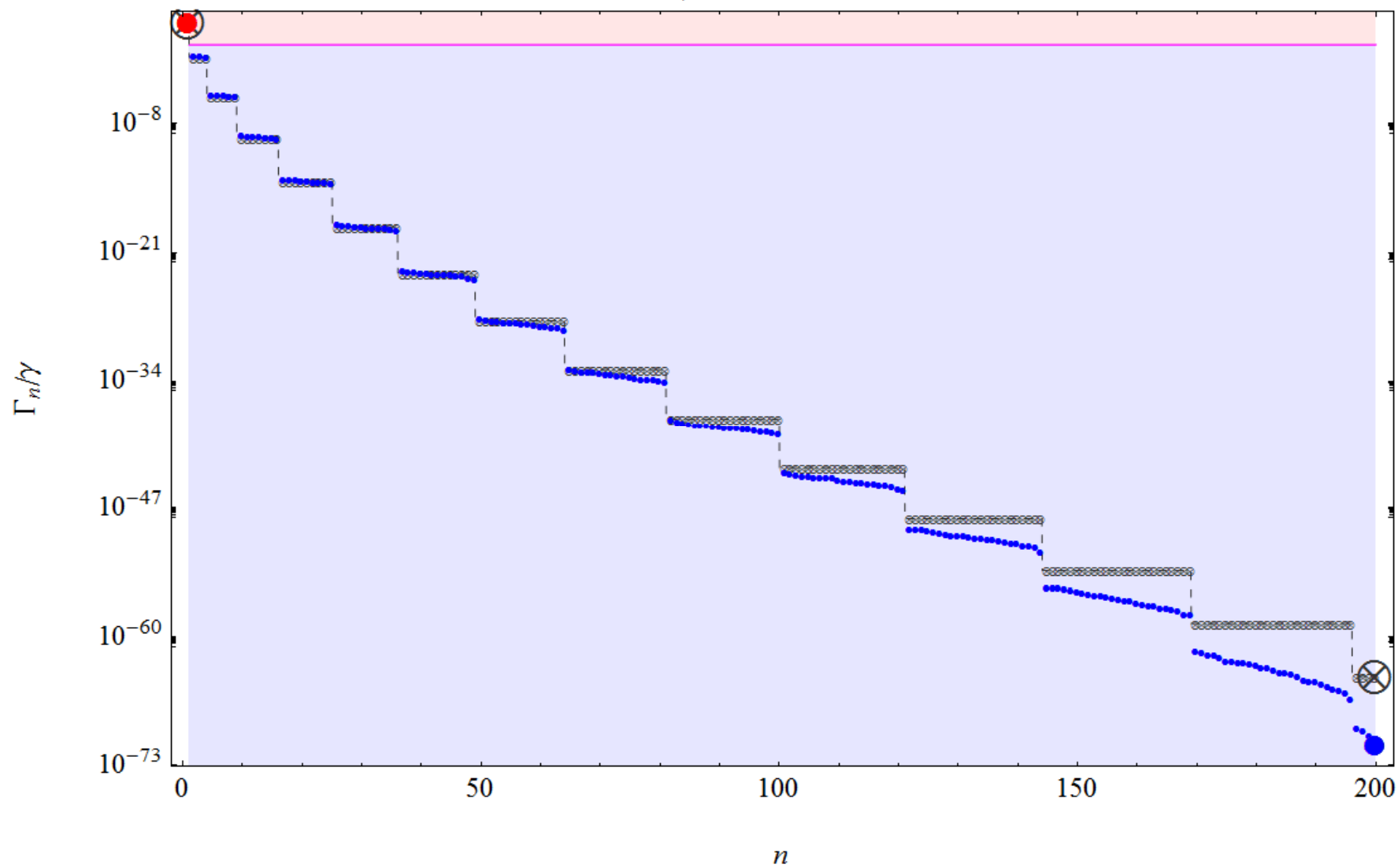
Sektion Physik der Universität München

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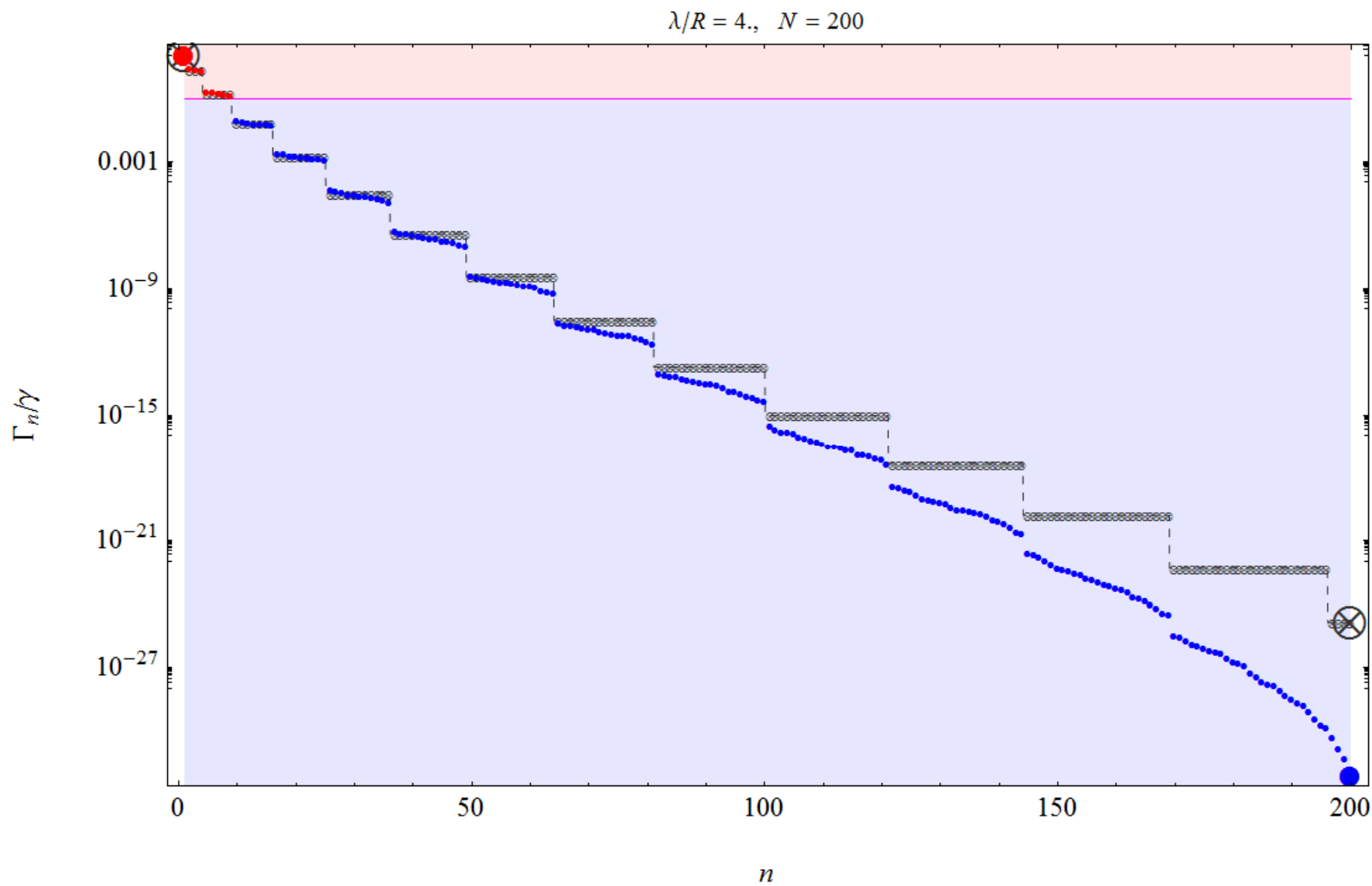
The theory of WEISSKOPF and WIGNER<sup>1</sup> of the natural line width is extended to a case where many atoms at given positions interact with one common quantized radiation field. At time  $t=0$ , one quantum of radiation energy is stored by the atoms, which however does not mean necessarily that exactly one of the atoms must be excited. We study the process of the creation of a photon in dependence on the positions of the atoms and on the state of the system of atoms at  $t=0$ . The dimensions of the emitted wave train and the decay time of this excited atomic state depend sensitively on these parameters.

# Large wavelength

$\lambda/R = 100.$ ,  $N = 200$

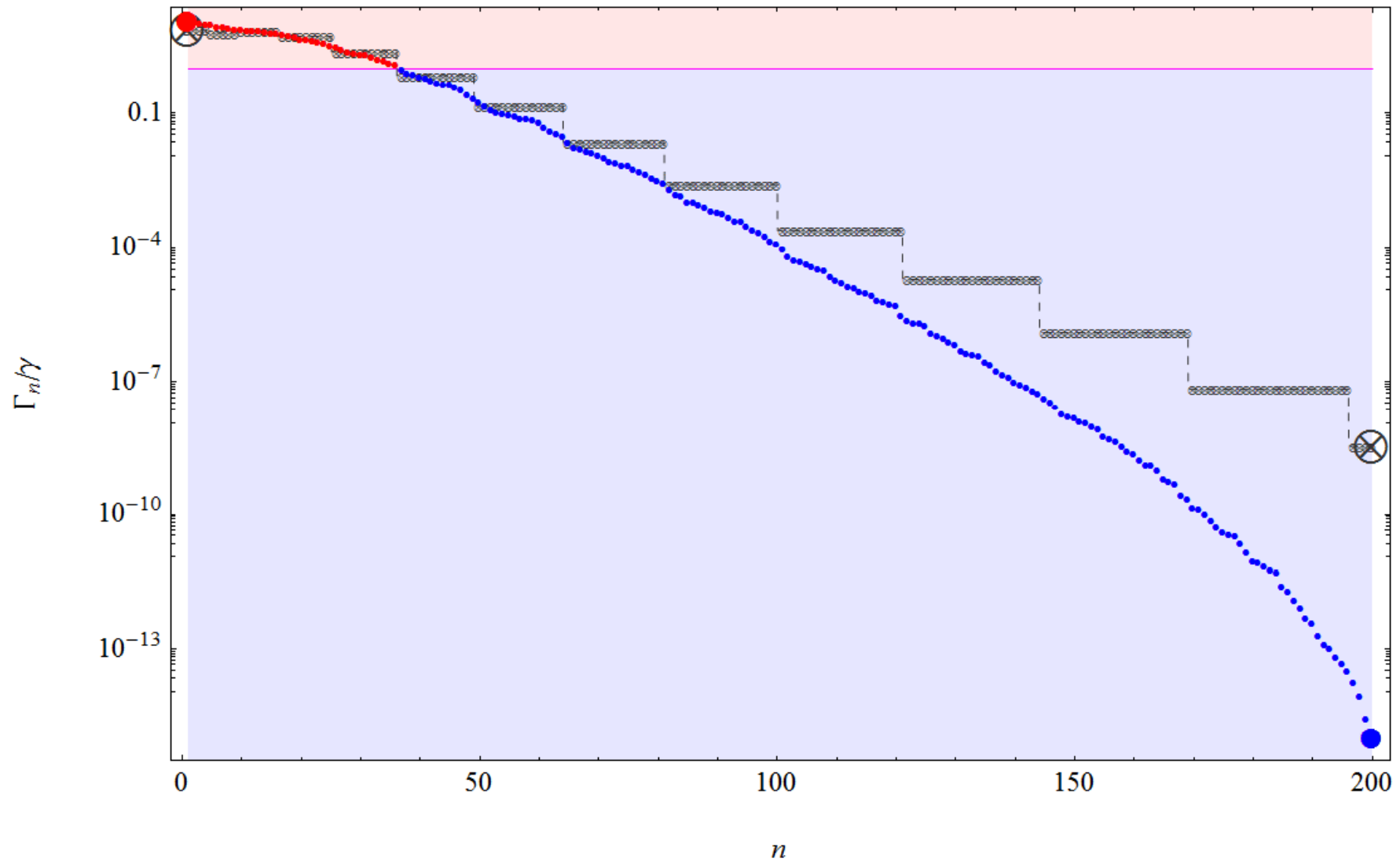


# Wavelength 4 times larger than size of sphere



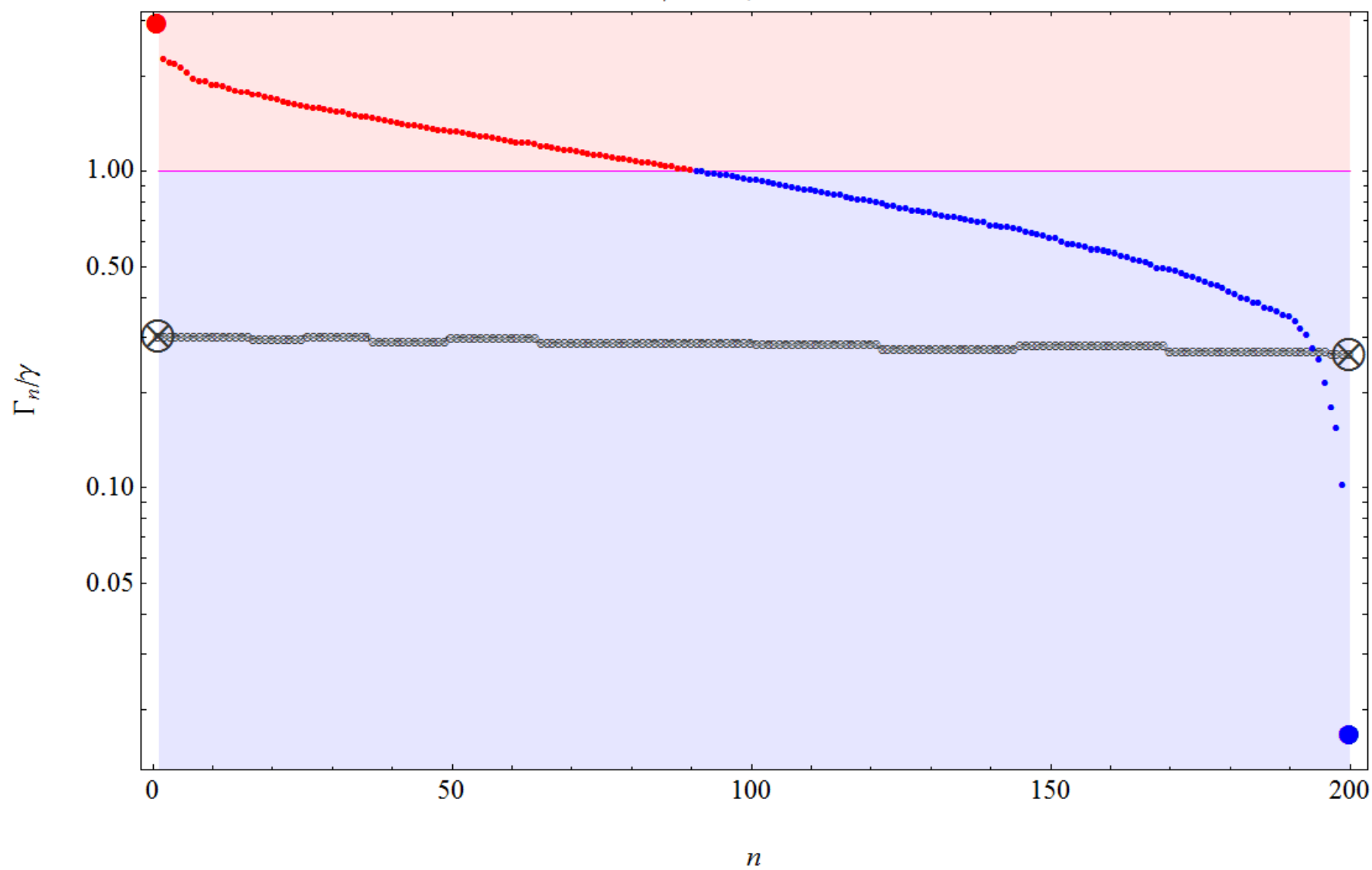
# Wavelength equal to radius of sphere

$\lambda/R = 1., N = 200$



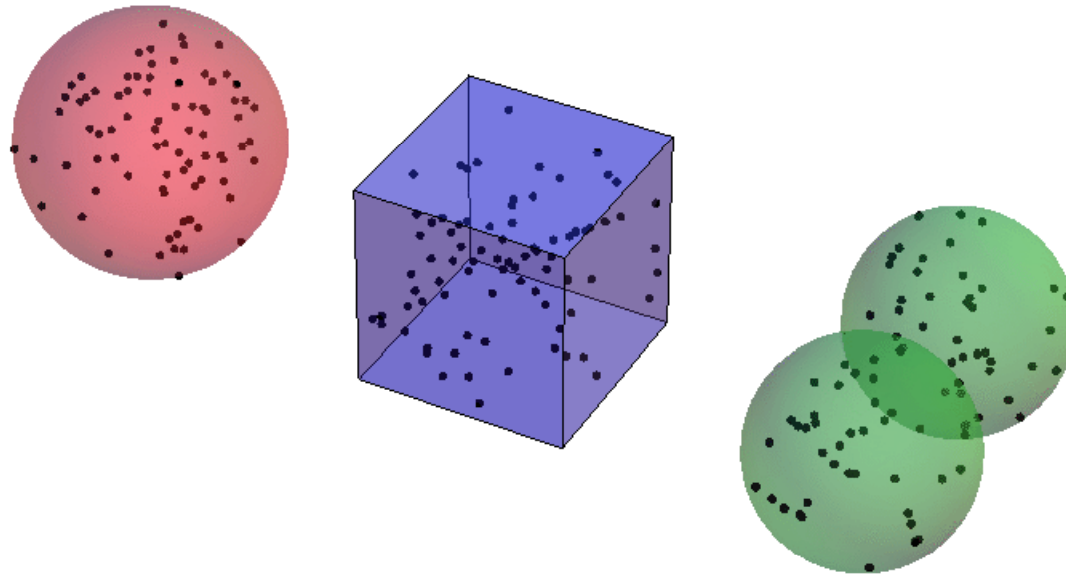
# Small wavelength

$\lambda/R = 0.2, N = 200$



## Dependence on shape

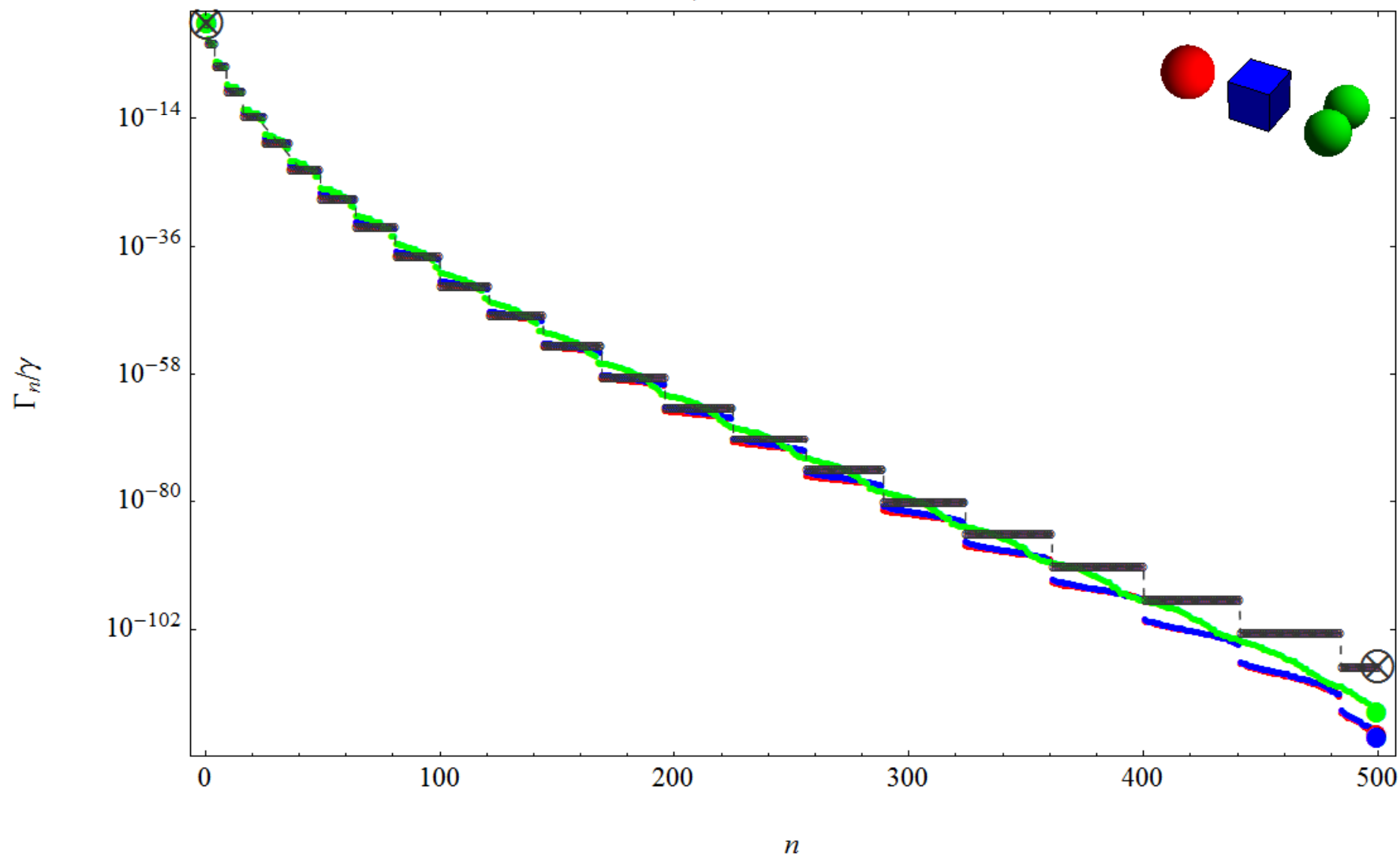
of linear dimensions large in comparison with  $\lambda/2$ . This means that  $\mathcal{C}_L^M(\mathbf{x})$  is still an approximate eigenstate for any large, three-dimensional region of a shape not too much different from a sphere, provided, of course, that the radiation law is isotropic. As for the latter, however, the





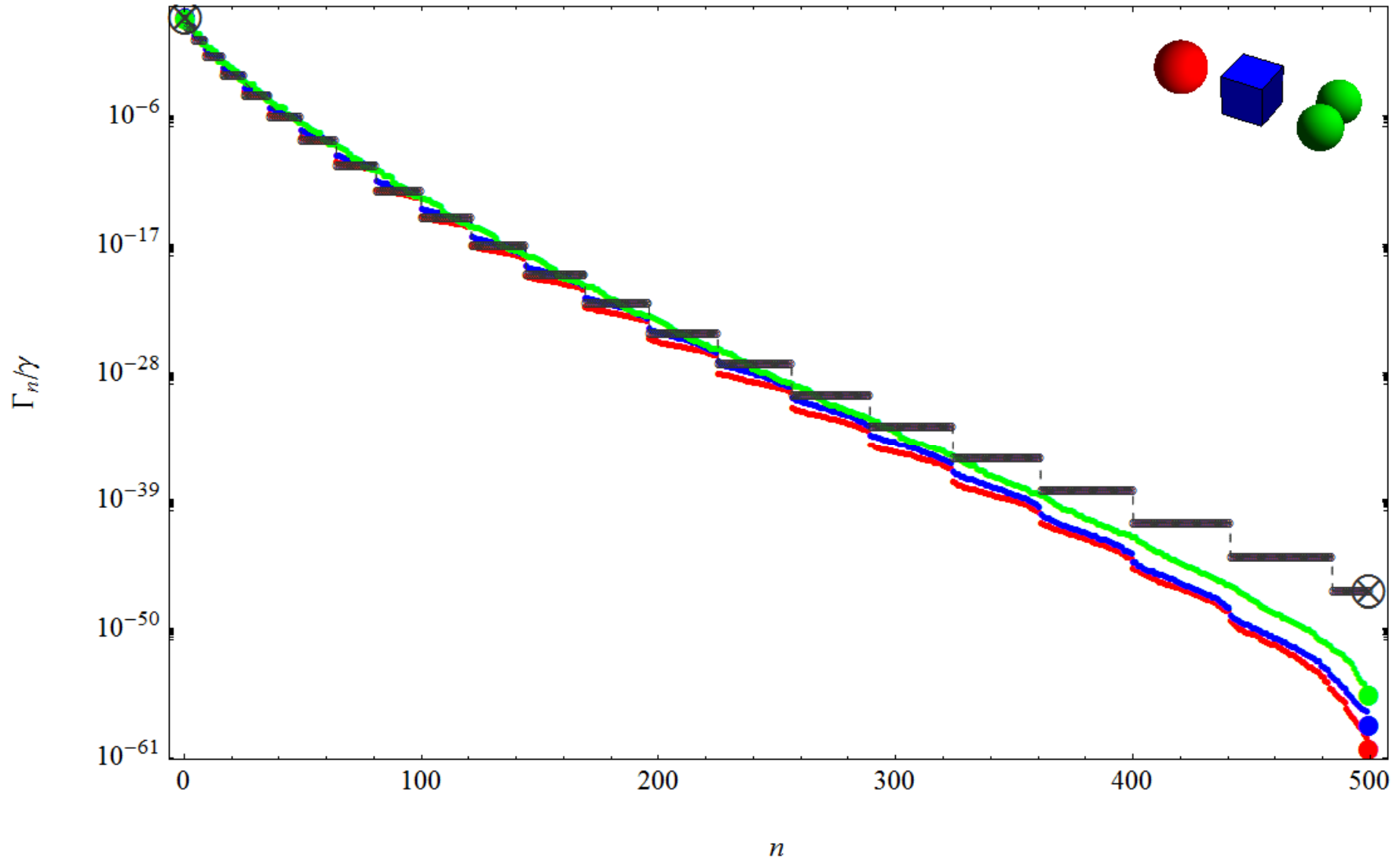
# Dependence on shape, large wavelength

$\lambda/R = 100.$ ,  $N = 500$

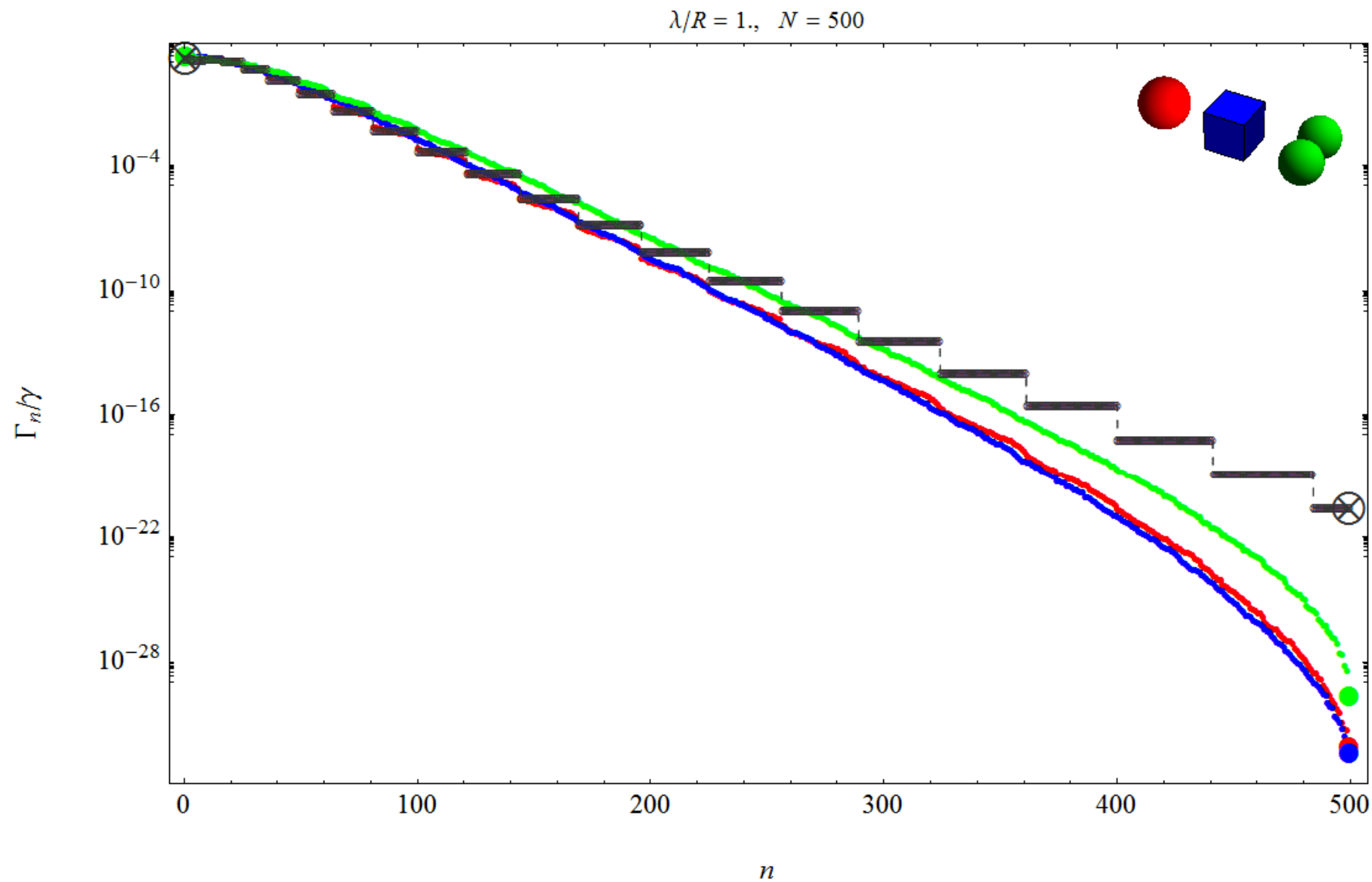


# Dependence on shape, wavelength larger than size

$\lambda/R = 4., N = 500$

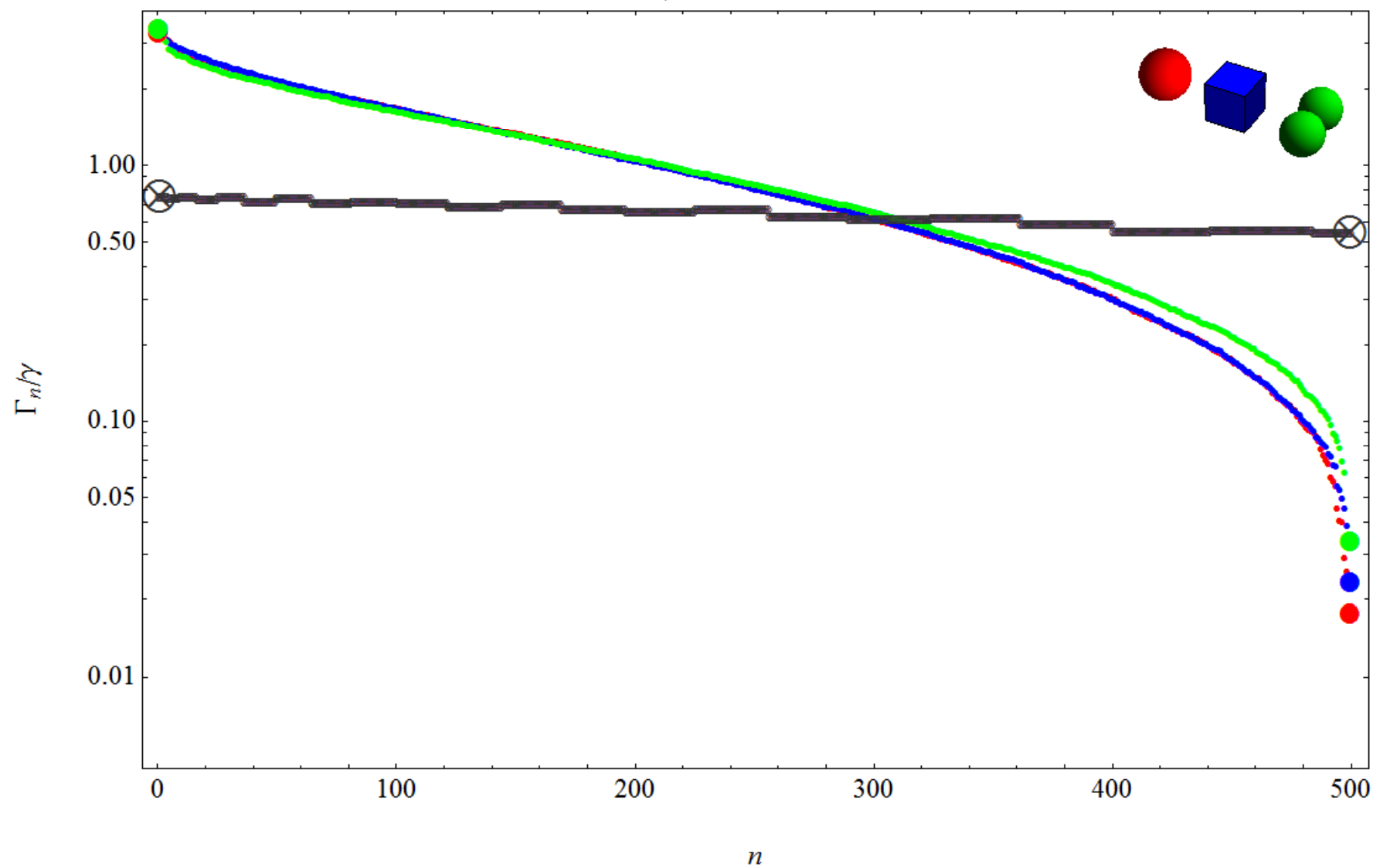


# Dependence on shape, wavelength comparable with size

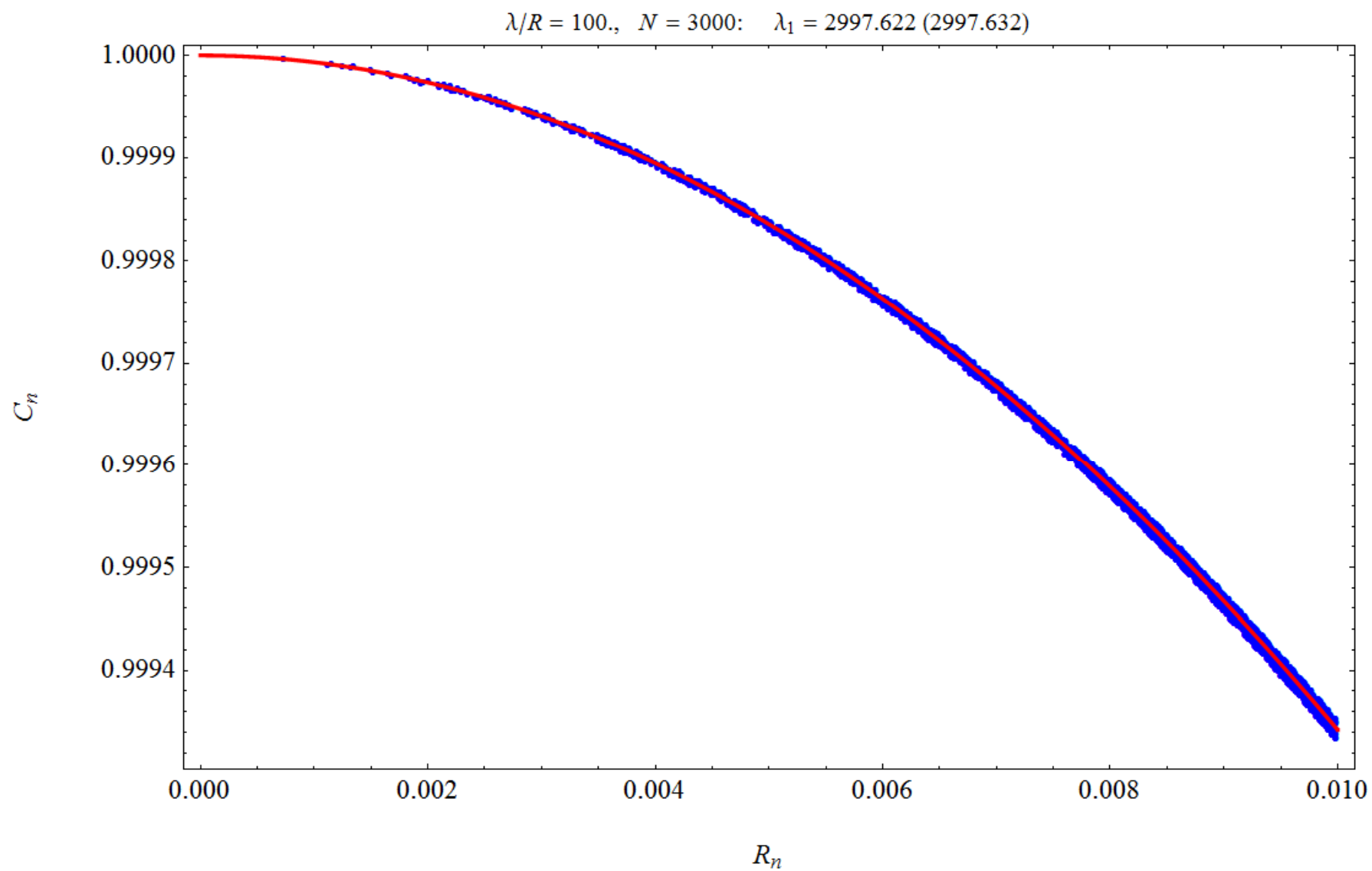


# Dependence on shape, short wavelength

$\lambda/R = 0.2, N = 500$

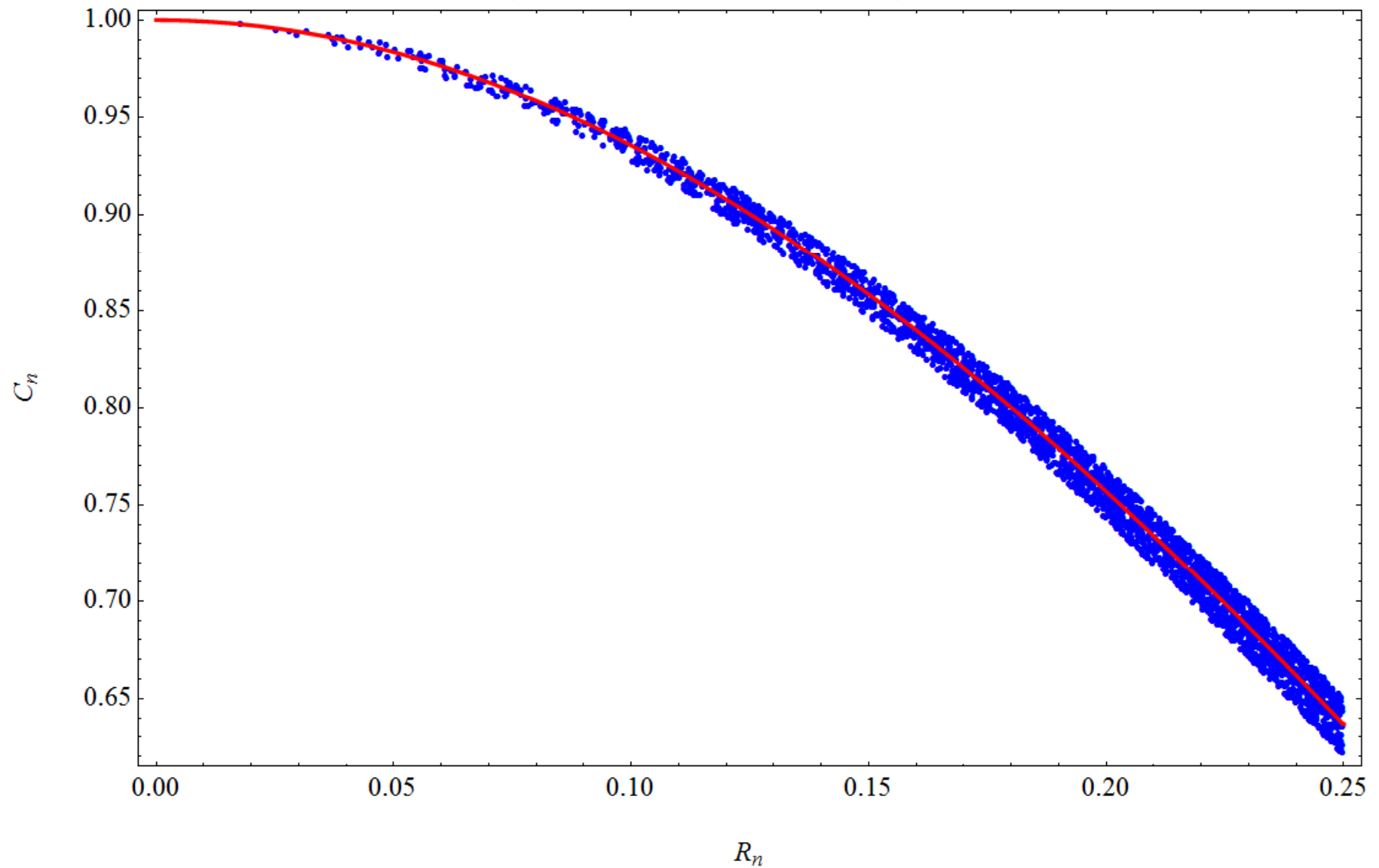


# Superradiant state vector, large wavelength



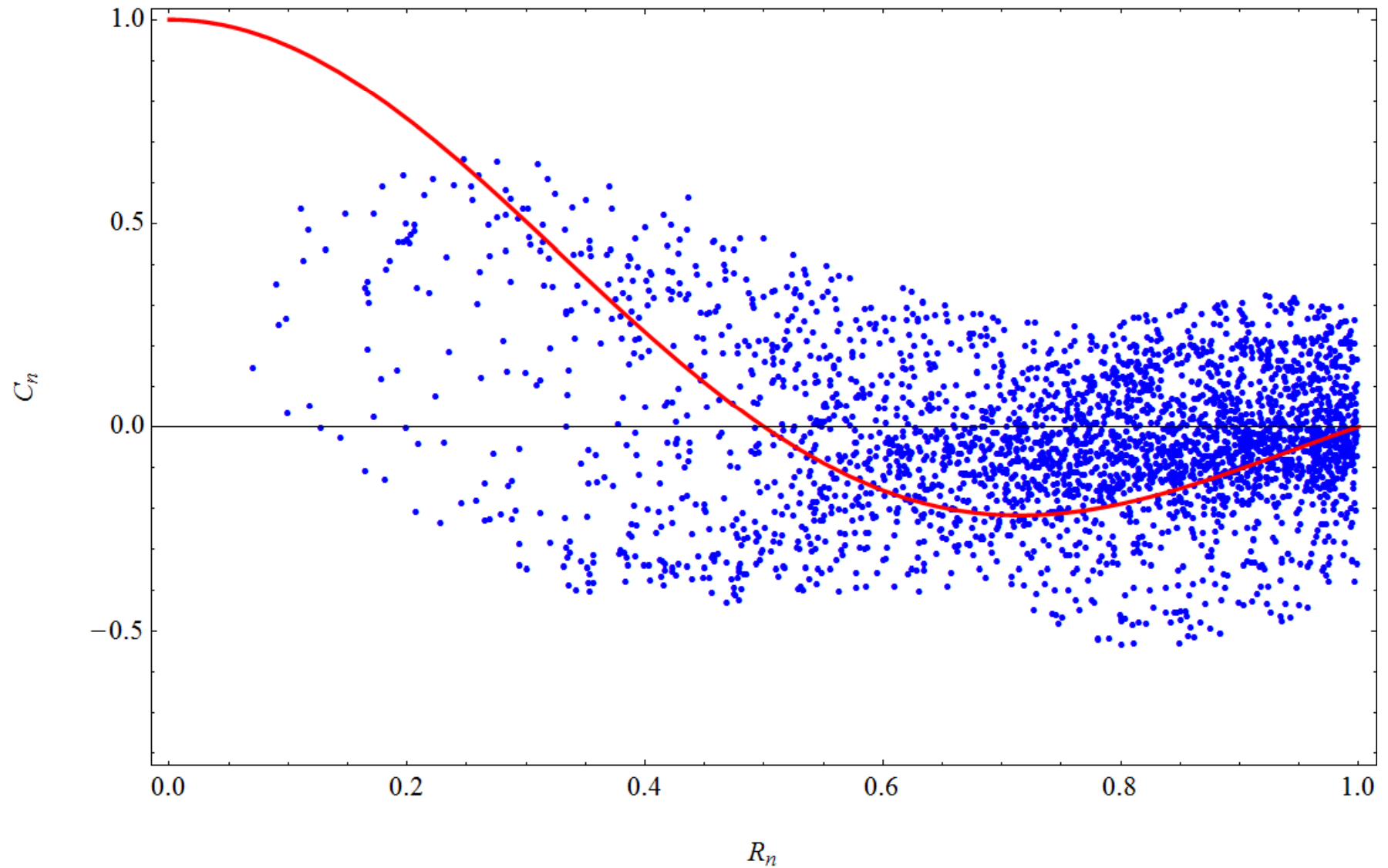
# Superradiant state vector, wavelength 4 times larger than size of sphere

$\lambda/R = 4.$ ,  $N = 3000$ :  $\lambda_1 = 1825.028$  (1823.781)



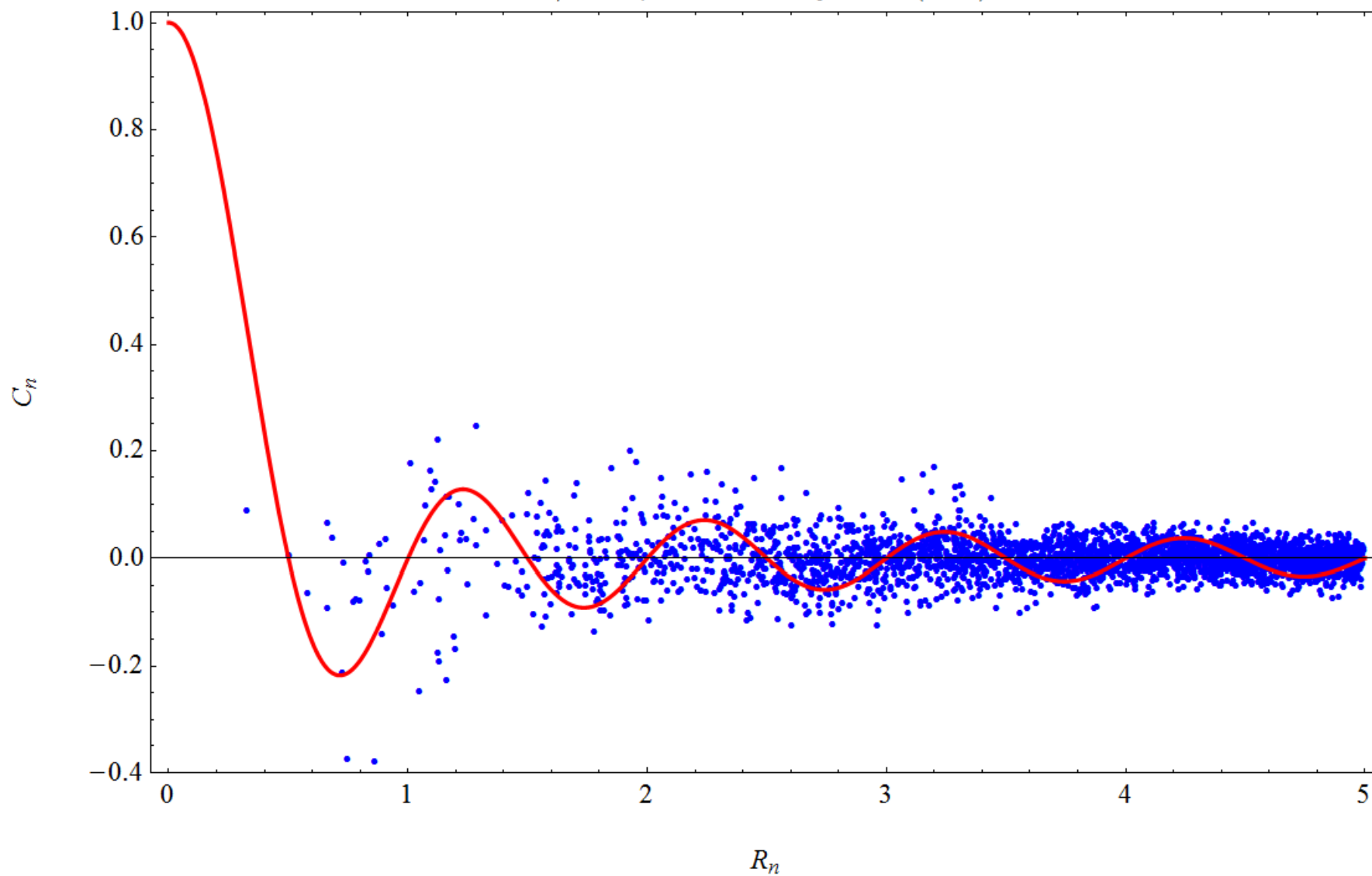
# Superradiant state vector, wavelength equal to radius of sphere

$\lambda/R = 1.$ ,  $N = 3000$ :  $\lambda_1 = 122.358$  (113.986)



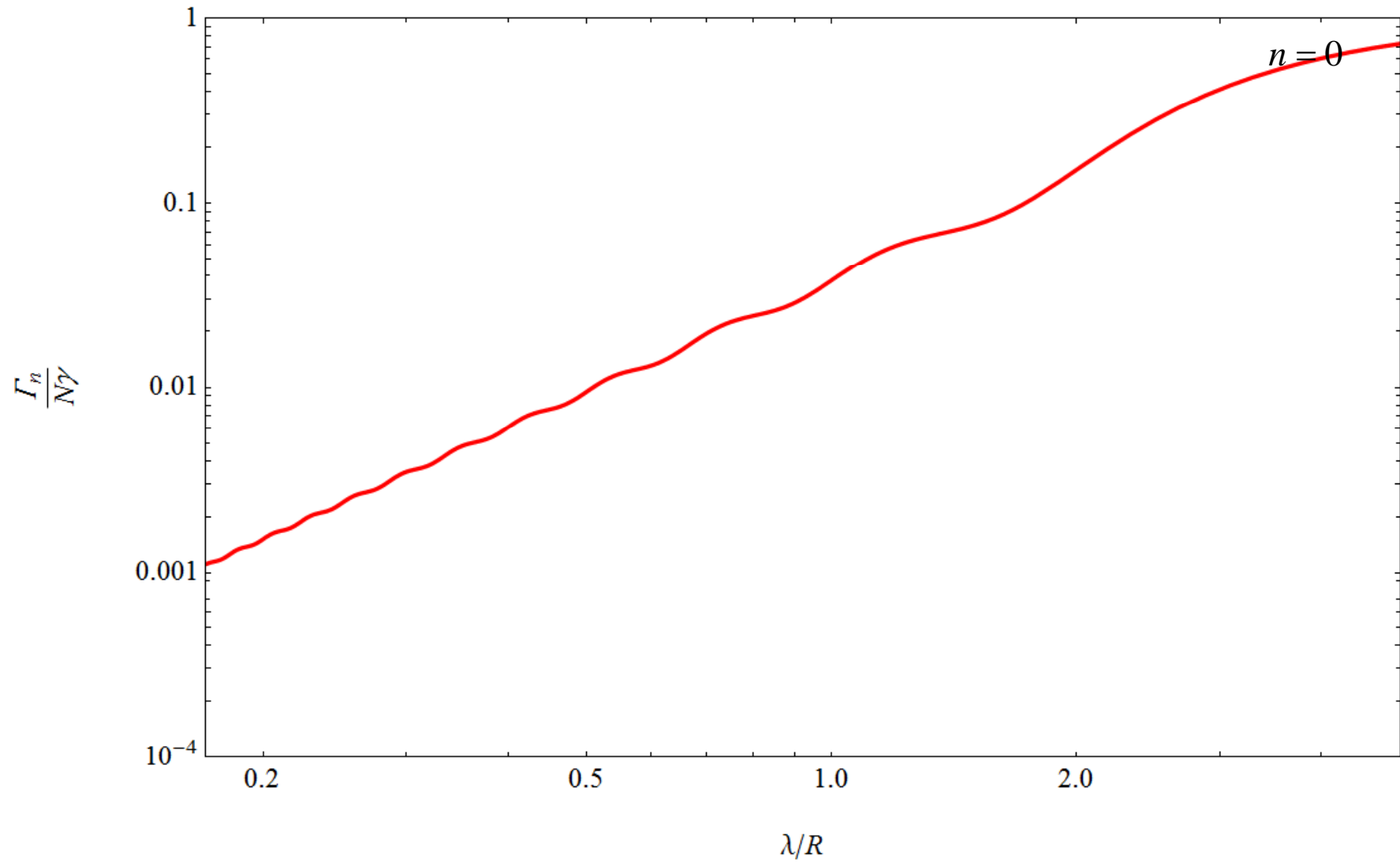
# Superradiant state vector, small wavelength

$\lambda/R = 0.2, N = 3000: \lambda_1 = 9.093 (4.559)$

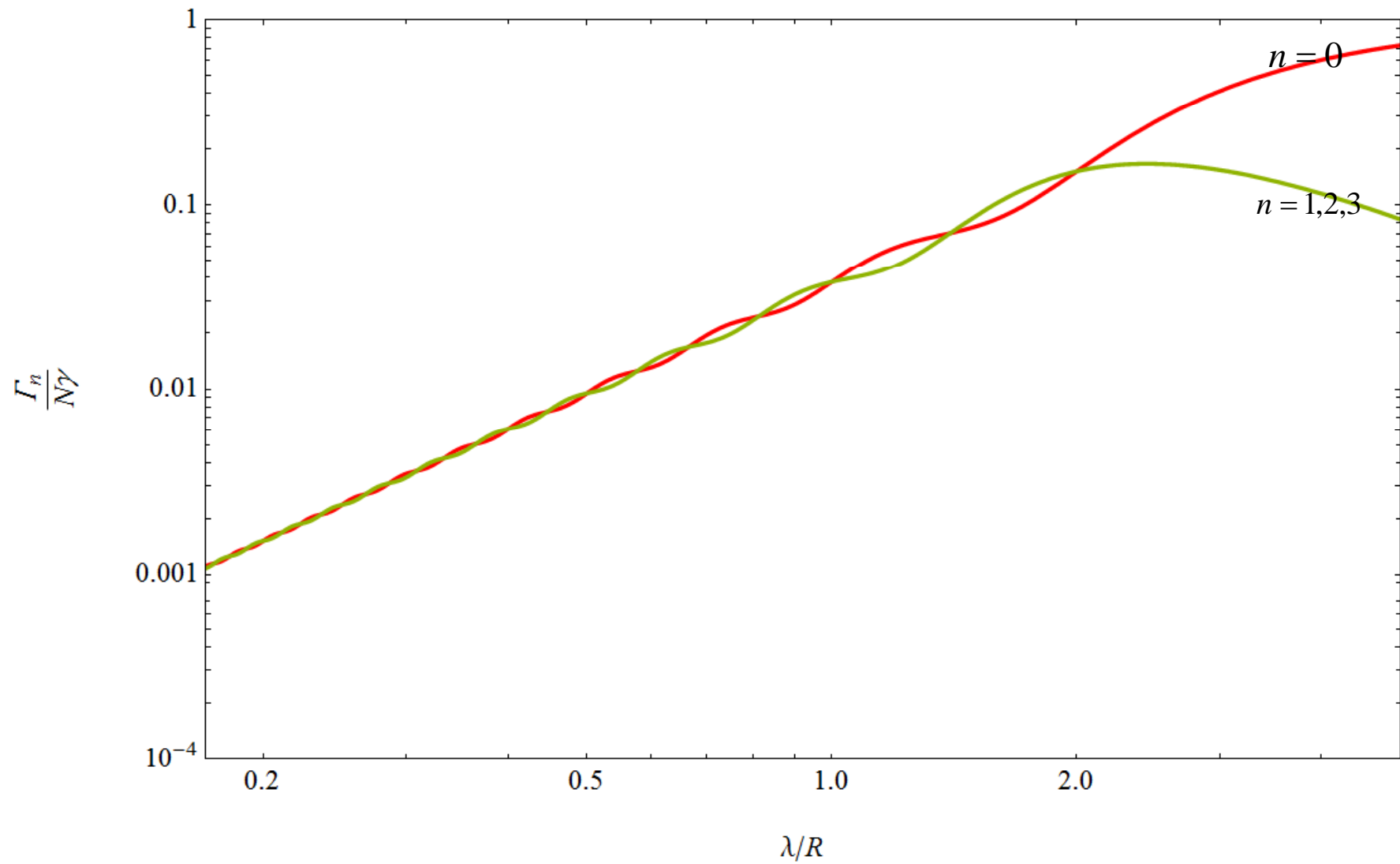




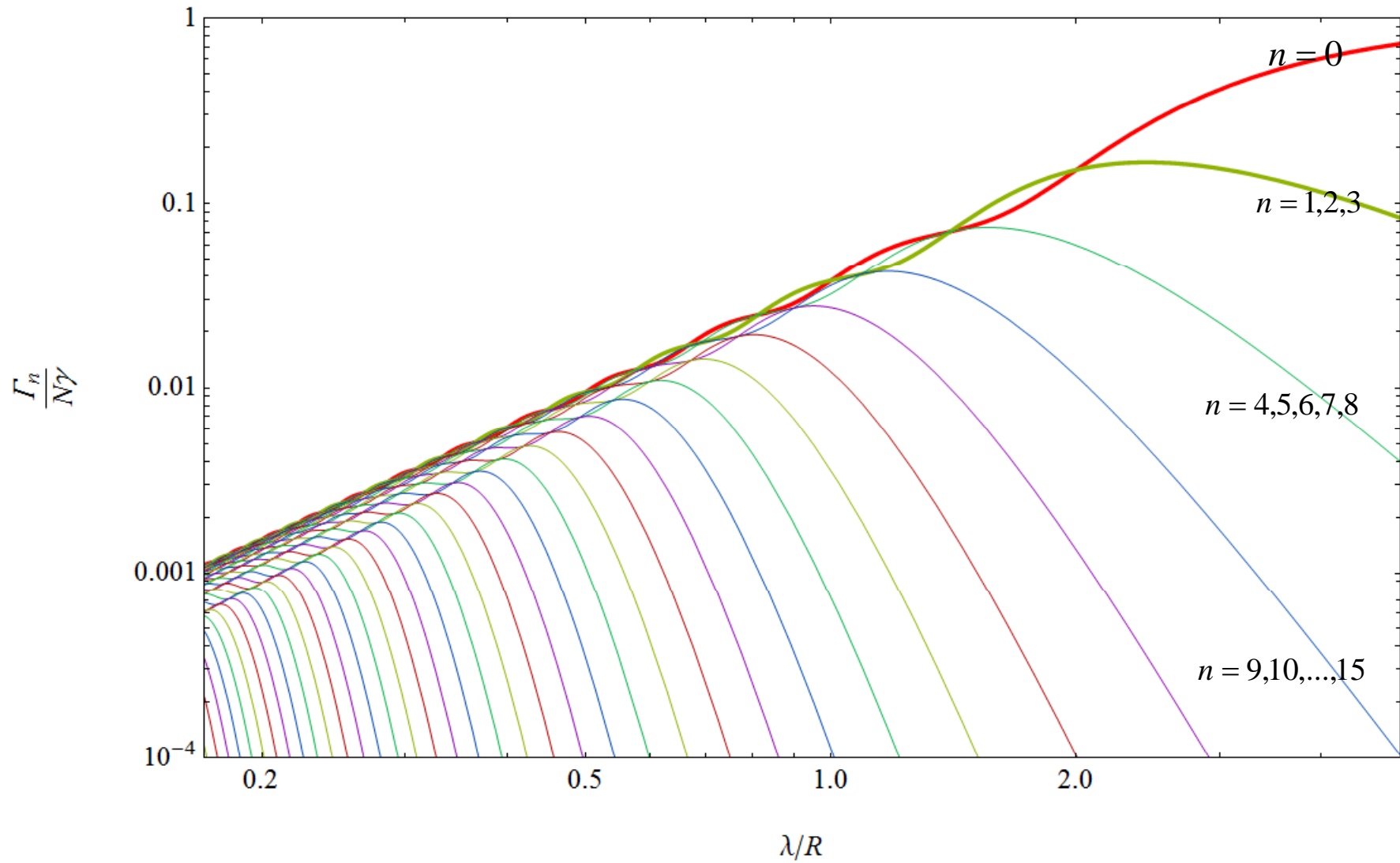
# Decay rates for different wavelength



# Decay rates for different wavelength



# Decay rates for different wavelength



# Calculating the decay rate of “superradiance” $\Gamma_0$

$$\Gamma_0^{(N \rightarrow \infty)} = \frac{3N}{2(k_0 R)^2} \left[ 1 - \frac{\sin(2k_0 R)}{2k_0 R} \right]$$

$$\boldsymbol{\beta}_{n,m} = \left\{ j_n(k_0 r_i) Y_{n,m}(\theta_i, \varphi_i) \right\}_{/i=1,2,3,\dots,N}, \quad \mathbf{B} = \left\{ \boldsymbol{\beta}_{n,m} \right\}_{/n \leq n_{\max}}$$

$$\mathbf{F}(n_{\max}) = \mathbf{BFB}^T, \quad \mathbf{S}(n_{\max}) = \mathbf{BB}^T$$

$\Gamma_0(n_{\max})$ : largest zero of polynomial  $|\mathbf{F}(n_{\max}) - \Gamma \mathbf{S}(n_{\max})|$ .

Size of the matrix :  $(n_{\max} - 1)^2$  instead of  $N$ .

# Numerical results for $\Gamma_0$

$N = 1000, \lambda / R = 1$

$\Gamma_0^{(N \rightarrow \infty)}$	38.00
$\Gamma_0(0)$	39.90
$\Gamma_0(1)$	44.05
$\Gamma_0(2)$	44.19
$\Gamma_0(3)$	45.96
$\Gamma_0(4)$	46.39
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$\Gamma_0^{(N)}$	46.55

$N = 1000, \lambda / R = 1/4$

$\Gamma_0^{(N \rightarrow \infty)}$	1.52
$\Gamma_0(0)$	2.48
$\Gamma_0(1)$	3.21
$\Gamma_0(2)$	3.28
...	...
$\Gamma_0(10)$	4.30
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$\Gamma_0^{(N)}$	4.63